

RKHS approach for signal detection in rotation and scale space random fields

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Two important papers of Worsley, Siegmund and coworkers consider rotation and scale space random fields for detecting signals in fMRI (functional magnetic resonance imaging) brain images. They use the global maxima of images for detection of a signal. In the current work, we utilize a reproducing kernel Hilbert space (RKHS) approach to show for both rotation and scale space random fields the global maximum of the image is indeed the likelihood ratio test statistic.

Keywords: Scale space random fields, rotation space random fields, reproducing kernel Hilbert space, likelihood ratio statistic.

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Introduction

For the past two decades, Gaussian random fields have been used in a variety of applications in astronomy, neural imaging and genetics to model images produced by modern sensor technologies. One of the main problems in these subject areas is searching for an activation or a signal in the images. Two papers of [1] and [2] consider rotation and scale space random fields for detecting signals in fMRI (functional magnetic resonance imaging) brain images. They use the global maxima of images for detection of a signal. For the scale space random field, [1] and for the rotation space random field, [2], without providing any proof, state that the global maximum, X_{\max} , is the likelihood ratio test static for testing the hypothesis of no signal. For the scale space random fields, [3] uses techniques of Gaussian measures on Hilbert spaces to present a proof. In the current work, we give another proof which works for both rotation and scale space random fields.

Reproducing kernel Hilbert spaces (RKHS) arise in different areas of statistics, including statistical signal processing, nonparametric curve estimation, random measures, and limit theorems. Use of the RKHS approach for the problem of signal detection in Gaussian noise now is a classical method in signal processing literature. [4, 5] used the RKHS approach to find the likelihood ratio for detection problem in the Gaussian case. These methods, for different situations, were extended by [6], [7], [8] among others. For a recent review see [9]. We utilize the RKHS theory to prove X_{\max} for both rotation and scale space random fields is indeed the likelihood ratio test statistic.

The rest of paper is divided into three sections. In Section 1, we introduce the scale and rotation space random fields. In Section 2, we review the main results from RKHS theory and its relation to the covariance function of a random field. Finally, in Section 3, we have the proof for X_{\max} being the likelihood ratio test statistic for testing no signal in rotation and scale space random fields.

1. Rotation and scale space random fields

In this section, we briefly introduce two Gaussian random fields extensively used for activation detection in fMRI images; Scale space random field and Rotation space random field.

1.1. Gaussian scale space random fields

Let k be an N -dimensional kernel such that

$$\int k^2(t)dt = 1.$$

[1] define the Gaussian scale space random field as

$$X(t, \sigma) = \sigma^{-N/2} \int k[\sigma^{-1}(h-t)] dZ(h), \quad (1.1)$$

where $Z(h)$ is a Gaussian random field defined on a subset of \mathbb{R}^N and σ is a positive constant.

A justification for working on Gaussian scale space random fields is as follows. Assume the random field $Z(t)$, $t \in \mathbb{R}^N$, satisfies

$$dZ(t) = \xi \sigma_0^{-N/2} k[\sigma_0^{-1}(t-t_0)] dt + dW(t), \quad (1.2)$$

where $t_0 \in T \subset \mathbb{R}^N$, $\xi \geq 0$ and $\sigma_0 > 0$ are fixed values and W is an N -dimensional Brownian sheet. The unknown parameter $\theta = (\xi, t_0, \sigma_0)$ represents the amplitude, location and scale of a signal added to the noise $dW(t)$. In other words, the shape of the signal in $dZ(t)$ matches the shape of the filter k . Models of this form have been used in different scientific contexts. See [2] and references therein. As [10] note, “from a statistical point of view, the scale space field is a continuous wavelet transform of white noise that is designed to be powerful at detecting a localized signal of unknown spatial scale and location.”

By replacing $dZ(h)$ from (1.2) in (1.1), the scale space random field takes the form of a signal plus noise model as

$$X(t, \sigma) = m(t, \sigma; \theta) + \varepsilon(t, \sigma), \quad (1.3)$$

where

$$m(t, \sigma; \theta) = (\sigma_0 \sigma)^{-1/2} \xi \int k[\sigma_0^{-1/2}(h-t_0)] k[\sigma^{-1/2}(h-t)] dh,$$

and

$$\varepsilon(t, \sigma) = \sigma^{-1/2} \int k[\sigma^{-1/2}(h-t)] dW(h)$$

is an $(N + 1)$ -dimensional zero mean Gaussian random field with the covariance function

$$R[(t_1, \sigma_1), (t_2, \sigma_2)] = (\sigma_1 \sigma_2)^{-N/2} \int k[\sigma_1^{-1}(h - t_1)] k[\sigma_2^{-1}(h - t_2)] dh.$$

See [1]. Note that $m(t, \sigma; \theta) = \xi R[(t, \sigma), (t_0, \sigma_0)]$.

For testing the hypothesis of no signal, that is, $\xi = 0$, [1] consider the test statistic that rejects for large positive values of

$$X_{\max} = \max_{(t, \sigma) \in T \times [\sigma_1, \sigma_2]} X(t, \sigma).$$

Then they use two different approaches, the expected Euler characteristic of the excursion set or the volume of tubes, to find an approximate P -value.

1.2. Rotation space random fields

A possible weakness of the scale space random field is a lack of power to detect signals that are not spherically symmetric. [2] generalize the scale space random field to the rotation space random field as follows.

A Gaussian rotation space random field is defined as

$$X(t, S) = \det(S)^{-\frac{1}{4}} \int k[S^{-\frac{1}{2}}(h - t)] dZ(t),$$

where k is spherically symmetric and S is a $N \times N$ positive definite matrix. The same argument as above for working on the scale space random field justifies working on the rotation space random field. Assume the random field $Z(t), t \in \mathbb{R}^N$, satisfies

$$dZ(t) = \xi \det(S_0)^{-\frac{1}{4}} k[S_0^{-\frac{1}{2}}(t - t_0)] dt + dW(t),$$

where S_0 is a member of a fixed parameter set Q of positive definite matrices. The unknown parameter $\theta = (\xi, t_0, S_0)$ represents amplitude, location, orientation and scale of signal and $dW(t)$ represents noise.

Similar to the scale space random field, the rotation space random field can be written as

$$X(t, S) = m(t, S; \theta) + \varepsilon(t, S), \quad (1.4)$$

where

$$m(t, S; \theta) = (\det(S_0 S))^{-\frac{1}{4}} \xi \int k[S_0^{-\frac{1}{2}}(h - t_0)] k[S^{-\frac{1}{2}}(h - t)] dh,$$

and

$$\varepsilon(t, S) = \det(S)^{-\frac{1}{4}} \int k[S^{-\frac{1}{2}}(h - t)] dW(h)$$

is an $(N + \frac{N(N+1)}{2})$ -dimensional zero mean Gaussian random field with the covariance function

$$R[(t_1, S_1), (t_2, S_2)] = \det(S_1 S_2)^{-\frac{1}{4}} \int k \left[S_1^{-\frac{1}{2}}(h - t_1) \right] k \left[S_2^{-\frac{1}{2}}(h - t_2) \right] dh. \quad (1.5)$$

See [2]. Again note that

$$m(t, S; \theta) = \xi R[(t, S), (t_0, S_0)]. \quad (1.6)$$

We use interchangeably $m(t, S; \theta)$ and $m(t, S)$ for the mean or the signal of the rotation space model.

For testing the hypothesis of no signal, that is, $\xi = 0$, [2] consider the test statistic that rejects for large positive values of

$$X_{\max} = \max_{(t, S) \in T \times Q} X(t, S).$$

Then they use two different approaches; the expected Euler characteristic of the excursion set or the volume of tubes, to find an approximate P -value for $N = 2$, when the kernel has a Gaussian form. [10] use the Gaussian kinematic formula to extend the P -value approximation for a general N and any kernel.

For the scale space random field [1] and for the rotation space random field [2], without providing any proof, state that X_{\max} is the likelihood ratio test static for testing the hypothesis of no signal. For the scale space random field, [11] use techniques of Gaussian measures on Hilbert spaces to present a proof. Here we use a reproducing kernel Hilbert space approach to prove X_{\max} is indeed the likelihood ratio test static for both rotation and scale space random fields.

2. Reproducing kernel Hilbert spaces and random fields

Reproducing kernel Hilbert spaces arise in different areas of statistics, including statistical signal processing, nonparametric curve estimation, random measures, and limit theorems. In this section, we briefly review the elements of RKHS theory and describe the RKHS generated by the covariance function of a random field.

Let \mathcal{H} be a Hilbert space of real valued functions defined on a set U with the inner product $\langle \cdot, \cdot \rangle_{\mathcal{H}}$ and norm $\| \cdot \|_{\mathcal{H}}$. The real valued function $K(x, y)$, where $x, y \in U$ is called the reproducing kernel of \mathcal{H} if the following conditions hold,

- i) for every $y \in U$, $K_y(x) = K(x, y)$ belongs to \mathcal{H} , as a function of x .
- ii) (reproducing property) for every $y \in U$ and $f \in \mathcal{H}$, $f(y) = \langle f, K_y \rangle_{\mathcal{H}}$.

The Hilbert space \mathcal{H} is called a RKHS if there exists a reproducing kernel for \mathcal{H} .

It's clear that a reproducing kernel is a positive semi definite function. The converse of this is a very important theorem for the utility of RKHS in the study of random fields.

Theorem 2.1. (Moore-Aronszajn Theorem) *Let U be an index set and $K : U \times U \rightarrow \mathbb{R}$ be a symmetric positive semi definite function. Then there exists a unique Hilbert space of functions on U for which K is a reproducing kernel.*

For a proof see [12].

Since the covariance function of a random field is a positive semi definite function by the Moore-Aronszajn Theorem, there exists a unique RKHS corresponding to a second order random field. In fact, this is a bridge from abstract spaces in mathematics to statistics and gives access to a powerful

tool for solving statistical signal detection problems. To see that, consider a second-order zero-mean random field $\{X(u), u \in U\}$ with covariance function K and let L_X^2 be the Hilbert space of random variables generated by X . Let \mathcal{H}_K denote the RKHS with reproducing kernel K . Consider the linear map J from L_X^2 to \mathcal{H}_K satisfying

$$J(X(u)) = K(u, \cdot), \quad u \in U. \quad (2.1)$$

The linear map J , known as the Loeve-Parzen canonical congruence, establishes an isomorphism between L_X^2 and \mathcal{H}_K as stated in the following theorem.

Theorem 2.2. *The linear map J is a one-to-one map and for all $\eta_1, \eta_2 \in L_X^2$ satisfies*

$$\langle \eta_1, \eta_2 \rangle_{L_X^2} = \langle J(\eta_1), J(\eta_2) \rangle_{\mathcal{H}_K}.$$

For a proof see [13].

The next step in using the RKHS approach for the problem of signal detection in Gaussian noise is to derive the density of two Gaussian measures with respect to each other if they are equivalent. [4, 5] used the RKHS approach to find the likelihood ratio for the detection problem in Gaussian case. These methods were extended for different situations by [6], [7], [8] among the others. For a recent review see [9]. Here we mention a result which suits our models.

Theorem 2.3. *Let U be a separable metric space and Ω be the set of real valued functions defined on U , and \mathcal{F} be the sigma field of cylinder sets. Consider a second-order zero-mean Gaussian random field $\{\varepsilon(u), u \in U\}$ with covariance function K and let*

$$X(u) = m(u) + \varepsilon(u).$$

Let P_X and P_ε be two Gaussian probability measures on (Ω, \mathcal{F}) induced by $X(u)$ and $\varepsilon(u)$, respectively. If $m(u) \in \mathcal{H}_K$, then P_X and P_ε are equivalent and

$$\frac{dP_X}{dP_\varepsilon}(x) = \exp \left[J^{-1}(m) - \frac{1}{2} \langle m, m \rangle_{\mathcal{H}_K} \right].$$

3. Likelihood ratio test for scale and rotation space random fields

In this section, we use the RKHS results to derive the likelihood ratio test statistics for scale and rotation space random fields. For this, what we need to do is to find

$$\lambda(x) = \max_{\theta^* \in \Theta} \frac{dP_{\theta^*}}{dP_{\theta_0}}(x), \quad (3.1)$$

where $\frac{dP_{\theta^*}}{dP_{\theta_0}}(x)$ is the Radon-Nikodym derivative of the measure under the alternative, P_{θ^*} , with respect to the measure P_{θ_0} under the null hypothesis of no signal, $\theta = \theta_0$. The likelihood ratio test rejects H_0 for small values of $1/\lambda(x)$ or equivalently for large values of $\lambda(x)$.

Theorem 3.1. Let $X(t, S)$ be the Gaussian rotation space random field defined in (1.4). For testing

$$\begin{cases} H_0 : X(t, S) = \varepsilon(t, S) \\ H_1 : X(t, S) = m(t, S) + \varepsilon(t, S) \end{cases}$$

the likelihood ratio test statistic is equivalent to

$$X_{\max} = \max_{(t, S) \in T \times Q} X(t, S).$$

Proof. From (1.6), $m(t, S) = \xi R[(t, S), (t_0, S_0)] \in \mathcal{H}_R$, where

$$R[(t_1, S_1), (t_2, S_2)] = \det(S_1 S_2)^{-1/4} \int k[S_1^{-1/2}(h - t_1)] k[S_2^{-1/2}(h - t_2)] dh.$$

To derive the likelihood ratio, by Theorem 2.3, we need to find $J^{-1}(m)$ and $\langle m, m \rangle_{\mathcal{H}_R} = \|m\|_{\mathcal{H}_R}$. By the very first definition (2.1) and linearity of J^{-1} , $J^{-1}(m) = \xi x(t_0, S_0)$. To find $\|m\|_{\mathcal{H}_R}$, note that from the reproducing property and (1.6) we have

$$\|m\|_{\mathcal{H}_R} = \xi^2 \|R[., (t_0, S_0)]\|_{\mathcal{H}_R}^2 = \xi^2 R[(t_0, S_0), (t_0, S_0)].$$

Since $R[(t_0, S_0), (t_0, S_0)] = \int k^2(h) dh = 1$, $\|m\|_{\mathcal{H}_R}^2 = \xi^2$. Therefore, the likelihood ratio $\lambda(x)$ becomes

$$\begin{aligned} \lambda(x) &= \exp \left[\xi x(t_0, S_0) - \frac{\xi^2}{2} \right] \\ &= \exp \left\{ -\frac{1}{2} [(\xi - x(t_0, S_0))^2 - x^2(t_0, S_0)] \right\}. \end{aligned}$$

Since the log function is a nondecreasing function, the likelihood ratio test statistic is equivalent to

$$\max_{(t_0, S_0, \xi) \in T \times Q \times (0, \infty)} \left\{ -\frac{1}{2} [(\xi - x(t_0, S_0))^2 - x^2(t_0, S_0)] \right\}$$

For every fixed value of (t_0, S_0) the maximum log likelihood occurs at $\xi = x(t_0, S_0)$ and since $\xi \geq 0$, we have

$$\hat{\xi} = \begin{cases} x(\hat{t}_0, \hat{S}_0) & \text{if } x(\hat{t}_0, \hat{S}_0) \geq 0 \\ 0 & \text{if } x(\hat{t}_0, \hat{S}_0) < 0 \end{cases},$$

where $x(\hat{t}_0, \hat{S}_0) = \max_{(t_0, S_0) \in T \times Q} x(t_0, S_0)$. Therefore, the likelihood ratio test statistic is equivalent to X_{\max} .

In a similar manner, we can prove the following theorem for the scale space random field.

Theorem 3.2. Let $X(t, \sigma)$ be the Gaussian scale space random field defined in (1.3). For testing $\xi = 0$ against $\xi > 0$, the likelihood ratio test statistic is equivalent to

$$X_{\max} = \max_{(t, \sigma) \in T \times [\sigma_1, \sigma_2]} X(t, \sigma).$$

Discussion

In this paper, we proved that for both rotation and scale space random fields the global maximum of the image is indeed the likelihood ratio test statistic. The result can be extended to some other random fields beyond these two random fields. As long as we know how to find $J^{-1}(m)$ and $\|m\|$, we can find the likelihood ratio test statistic. It is not very obvious to the authors how to find $J^{-1}(m)$ and $\|m\|$ for any given random field. But it is not very hard to extend the result to the case that the random field $Z(t)$, for $t \in \mathbb{R}^N$, defined as

$$dZ(t) = \sum_{l=1}^L \xi_l \sigma_{0l}^{-N/2} k[\sigma_{0l}^{-1}(t - t_{l0})] dt + dW(t), \quad (3.2)$$

where $t_{l0} \in T \subset \mathbb{R}^N$, ξ_l , and $\sigma_{0l} > 0, l = 1, \dots, L$ are fixed values and W is an N -dimensional Brownian sheet. The unknown parameter $\theta_l = (\xi_l, t_{l0}, \sigma_{0l})$ represents the amplitude, location and scale of the l -th signal added to the noise $dW(t)$.

This case and some other cases along with the approximate distribution of the likelihood ratio test statistic is an undergoing work and will be presented in a future publication.

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